

# Physics 491 - Quantum Mechanics I

## Problem Set #4 - Solutions

### Prob. 1: Show + Answers

(a) Part 1: Linear Operators:

An operator  $\hat{O}: f(x)$  is linear iff

$$\hat{O}: (c_1 f_1(x) + c_2 f_2(x)) = c_1 \hat{O}: f_1(x) + c_2 \hat{O}: f_2(x)$$

•  $\hat{O}_1: \psi(x) \equiv x^3 \psi(x)$

$$\begin{aligned} \hat{O}_1: (c_1 \psi_1(x) + c_2 \psi_2(x)) &= x^3 (c_1 \psi_1(x) + c_2 \psi_2(x)) \\ &= c_1 (x^3 \psi_1(x)) + c_2 (x^3 \psi_2(x)) = c_1 \hat{O}_1 \psi_1(x) + c_2 \hat{O}_1 \psi_2(x) \end{aligned}$$

$$\Rightarrow \boxed{\hat{O}_1 \text{ is a linear operator}}$$

•  $\hat{O}_2: \psi(x) \equiv x \frac{d\psi}{dx}$

$$\begin{aligned} \hat{O}_2: (c_1 \psi_1(x) + c_2 \psi_2(x)) &= x \frac{d}{dx} (c_1 \psi_1(x) + c_2 \psi_2(x)) \\ &= c_1 \left( x \frac{d\psi_1}{dx} \right) + c_2 \left( x \frac{d\psi_2}{dx} \right) = c_1 \hat{O}_2 \psi_1(x) + c_2 \hat{O}_2 \psi_2(x) \end{aligned}$$

$$\Rightarrow \boxed{\hat{O}_2 \text{ is a linear operator}}$$

- $\sigma_3: \psi(x) \equiv \lambda \psi^*(x)$

$$\begin{aligned} \sigma_3: (c_1 \psi_1(x) + c_2 \psi_2(x)) &= \lambda (c_1^* \psi_1^*(x) + c_2^* \psi_2^*(x)) \\ &= c_1^* (\lambda \psi_1^*(x)) + c_2^* (\lambda \psi_2^*(x)) = c_1^* \sigma_3: \psi_1(x) + c_2^* \sigma_3: \psi_2(x) \\ &\neq c_1 \sigma_3: \psi_1 + c_2 \sigma_3: \psi_2 \end{aligned}$$

$\Rightarrow$   $\sigma_3$  is not a linear operator

- $\sigma_4: \psi(x) \equiv e^{\psi(x)}$

$$\begin{aligned} \sigma_4: (c_1 \psi_1(x) + c_2 \psi_2(x)) &= e^{c_1 \psi_1(x) + c_2 \psi_2(x)} \\ &= e^{c_1 \psi_1(x)} e^{c_2 \psi_2(x)} \neq c_1 \sigma_4: \psi_1 + c_2 \sigma_4: \psi_2 \end{aligned}$$

$\Rightarrow$   $\sigma_4$  is not a linear operator

- $\sigma_5: \psi(x) \equiv \frac{d\psi}{dx}$

$$\begin{aligned} \sigma_5: (c_1 \psi_1(x) + c_2 \psi_2(x)) &= \frac{d}{dx} (c_1 \psi_1(x) + c_2 \psi_2(x)) \\ &= c_1 \frac{d\psi_1}{dx} + c_2 \frac{d\psi_2}{dx} = c_1 \sigma_5: \psi_1 + c_2 \sigma_5: \psi_2 \end{aligned}$$

$\Rightarrow$   $\sigma_5$  is a linear operator

$$\hat{O}_6: \psi(x) = \int_{-\infty}^x dx' (\psi(x') x')$$

$$\hat{O}_6: (c_1 \psi_1(x) + c_2 \psi_2(x)) = \int_{-\infty}^x dx' ([c_1 \psi_1(x') + c_2 \psi_2(x')] x')$$

$$= c_1 \int_{-\infty}^x dx' \psi_1(x') x' + c_2 \int_{-\infty}^x dx' \psi_2(x') x'$$

$$= c_1 \hat{O}_6: \psi_1(x) + c_2 \hat{O}_6: \psi_2(x)$$

$$\Rightarrow \boxed{\hat{O}_6 \text{ is a linear operator}}$$

(b)

Electron, mass  $m = 0.9 \times 10^{-30} \text{ kg}$  in an infinitely high box, dimension  $a = 10^{-9} \text{ m}$

The energy levels:  $E_n = n^2 \frac{\pi^2 \hbar^2}{2ma^2} = n^2 \frac{\pi^2 (\hbar c)^2}{2(mc^2)a^2}$

Useful units  $\hbar c = 1974 \text{ eV} \cdot \text{\AA}$  (check this!)

Electron mass  $mc^2 = 0.5 \text{ MeV} = 500 \text{ eV}$

$a = 10^{-9} \text{ m} = 10 \text{ \AA}$

$$\Rightarrow E_n = n^2 (0.38 \text{ eV})$$

• Energy difference  $\boxed{E_2 - E_1 = (4 - 1) 0.38 \text{ eV} = 1.15 \text{ eV}}$

• According to Bohr, the frequency of photon emitted:  $\omega = \frac{E_2 - E_1}{\hbar} = \frac{2\pi c}{\lambda} \Rightarrow \lambda = \frac{hc}{E_2 - E_1}$

$$\Rightarrow \boxed{\lambda = 1.07 \mu\text{m}}$$

## Problem 2

(a) Consider a general wave packet in free space at  $t=0$ :

$$(1) \quad \Psi(x,0) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} b(k) e^{ikx} dk = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \phi(k;x) dk$$

i.e. an infinite superposition of plane waves (free-space stationary states) with arbitrary coefficients,  $b(k)$ . Including the coefficients  $b(k)$ ;

$$(2) \quad b(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \Psi(x,0) e^{-ikx} dx \quad (\text{i.e. Fourier transform})$$

in the stationary state  $\phi(k;x)$  gives the last integral in (1) above

But we know the time dependence of the "stationary states"  $\phi(k;x)$  from the TDSE:

$$i\hbar \frac{\partial}{\partial t} \phi(k;x,t) = \hat{H} \phi(k;x,t) \quad \text{with } \hat{H} = \frac{\hat{p}^2}{2m} \text{ for free space. } \Rightarrow$$

$$(3) \quad \phi(k;x,t) = \phi(k;x,0) e^{-i\omega t}$$

where  $\omega \equiv \omega(k)$  is given by  $E_k = \hbar\omega(k) = \frac{\hbar^2 k^2}{2m}$ .  $\left(\frac{p^2}{2m}\right)$

Since we know the time dependence of the individual terms in the integral, we just integrate to get the overall time dependence which is then not stationary.

$$(4) \quad \Psi(x,t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \phi(k;x,t) dk = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} b(k) e^{ikx - \omega(k)t} dk$$

using (2), being careful with  $x$  and  $x'$  in (4)

$$b(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \psi(x', 0) e^{-ikx'} dx'$$

$$\psi(x, t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dk \left\{ \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} dx' \psi(x', 0) e^{-ikx'} dx' \right\} e^{i(kx - \omega t)}$$

$$= \int_{-\infty}^{\infty} dx' \psi(x', 0) \underbrace{\int_{-\infty}^{\infty} \frac{dk}{2\pi} e^{-i[\omega t - k(x-x')]}}_K$$

Evaluate second integral by completing square with  $2\alpha = \frac{\hbar k^2}{2m}$

$$A^2 k^2 + 2ABK(x-x') + B^2(x-x')^2 = \frac{\hbar t}{2m} k^2 - k(x-x') + B^2(x-x')^2$$

$$A = \sqrt{\frac{\hbar t}{2m}}$$

$$2AB = -1$$

$$B = -\frac{1}{2} \sqrt{\frac{2m}{\hbar t}}$$

$$B^2 = \frac{m}{2\hbar t}$$

$$K = \int_{-\infty}^{\infty} \frac{dk}{2\pi} e^{-i \left[ \sqrt{\frac{\hbar t}{2m}} k - \sqrt{\frac{m}{2\hbar t}} (x-x') \right]^2} e^{i \frac{m}{2\hbar t} (x-x')^2}$$

$$\text{let } u = \sqrt{\frac{\hbar t}{2m}} k - \sqrt{\frac{m}{2\hbar t}} (x-x')$$

$$du = \sqrt{\frac{\hbar t}{2m}} dk$$

$$K = \frac{1}{2\pi} \sqrt{\frac{2m}{\hbar t}} \int_{-\infty}^{\infty} du e^{-iu^2} \cdot e^{i \frac{m}{2\hbar t} (x-x')^2} = \frac{1}{2\pi} \sqrt{\frac{2\pi m}{i\hbar t}} e^{i \frac{m}{2\hbar t} (x-x')^2}$$

$\sqrt{\frac{2\pi}{i}}$  from Schenck's Math. Handbook

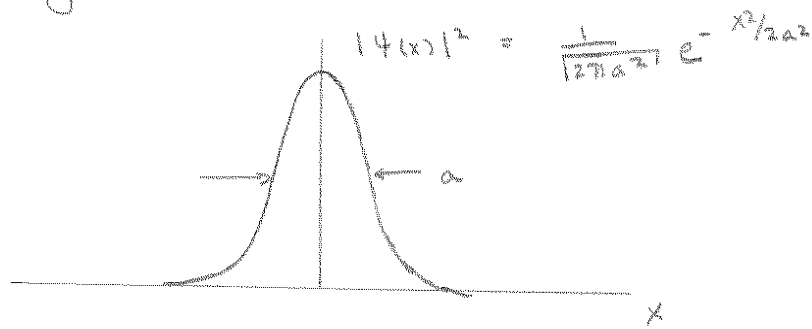
$$K = \sqrt{\frac{m}{2\pi i \hbar t}} e^{i \frac{m}{2\hbar t} (x-x')^2}$$

$$= K(x, x', t)$$

FREE SPACE PROPAGATOR

$$\psi(x, t) = \int_{-\infty}^{\infty} dx' \psi(x', 0) K(x, x', t)$$

(b) The probability distribution looks like:



The width  $a$  gives us an estimate of  $\Delta x$  at  $t=0$ .  
 Since this is the only length in the problem, we take it as our "characteristic" length,  $x_c$ .

$$x_c \equiv a \approx \Delta x$$

The Heisenberg uncertainty principle gives:

$$\Delta x \Delta p \geq \hbar$$

So taking the minimum momentum uncertainty at  $t=0$  for our characteristic momentum:

$$p_c = \Delta p_{\min}(t=0) = \frac{\hbar}{\Delta x(t=0)} = \frac{\hbar}{x_c} = \frac{\hbar}{a}$$

And energy: 
$$E_c = \frac{p_c^2}{2m} = \frac{\hbar^2}{2ma^2}$$

Associated with any characteristic energy is a characteristic time  $t_c \sim \frac{\hbar}{E_c} = \frac{2ma^2}{\hbar}$ .

Summarizing: 
$$x_c = a \quad p_c = \frac{\hbar}{a} \quad E_c = \frac{\hbar^2}{2ma^2} \quad t_c = \frac{2ma^2}{\hbar}$$

For  $m = 1g$ ,  $a = 1cm$ :

$$t_c = \frac{2(1g)(1cm^2)}{1.05 \times 10^{34} \text{ J}\cdot\text{s}} = \frac{2}{1.05} \frac{10^{-3} \frac{kg}{g} \cdot 10^{-4} \frac{m^2}{cm^2}}{10^{34} \frac{kg \cdot m^2}{g \cdot s}}$$

$$\approx 2 \times \frac{10^{-7}}{10^{34}} \text{ s}$$

$$t_c \approx 2 \times 10^{-27} \text{ s}$$

much larger than the age of the universe, so we don't expect to see spreading of objects anywhere near macroscopic size.

(c) Given  $\psi(x,0) = \frac{1}{(2\pi a^2)^{1/4}} e^{ik_0 x} e^{-\frac{x^2}{4a^2}}$

$$\psi(x,t) = \int_{-\infty}^{\infty} dx' K(x,x';t) \psi(x',0), \quad K(x,x') = \sqrt{\frac{m}{2\pi i \hbar t}} \exp\left[\frac{im}{2\hbar t} (x-x')^2\right]$$

Let us define  $t_c = \frac{2ma^2}{\hbar}$  (characteristic time scale, up to factor of 2)

$$\Rightarrow K(x,x';t) = \frac{1}{\sqrt{4\pi a^2}} \sqrt{\frac{t_c}{it}} \exp\left[i \frac{(x-x')^2}{4a^2} \frac{t_c}{t}\right]$$

$$\Rightarrow \psi(x,t) = \frac{1}{(2\pi a^2)^{1/4}} \frac{1}{\sqrt{4\pi a^2}} \sqrt{\frac{t_c}{it}} \int_{-\infty}^{\infty} dx' \underbrace{\exp\left[i \frac{(x-x')^2}{4a^2} \frac{t_c}{t}\right] \exp\left[-\frac{x'^2}{4a^2} + ik_0 x'\right]}_{\equiv \mathcal{I}}$$

$$\mathcal{I} = e^{i \frac{x^2}{4a^2} \frac{t_c}{t}} \exp\left\{ \underbrace{-\frac{1}{4a^2} \left(1 - i \frac{t_c}{t}\right) x'^2}_{=A} - 2 \underbrace{\left[ \frac{i}{4a^2} \left(\frac{t_c}{t}\right) x - \frac{ik_0}{2} \right] x'}_{=B} \right\}$$

$$= \exp\left[-A \left(x' - \frac{B}{A}\right)^2\right] e^{-\frac{B^2}{A}}$$

Aside  $\int_{-\infty}^{\infty} dx' e^{-A(x' - \frac{B}{A})^2} = \sqrt{\frac{\pi}{A}}$  when  $\text{Re}(A) \geq 0$

$$\Rightarrow \psi(x,t) = \frac{1}{(2\pi a^2)^{1/4}} \sqrt{\frac{t_c}{it(1-i\frac{t_c}{t})}} \exp\left\{ \frac{-\frac{1}{4a^2} \left[ \left(\frac{t_c}{t}\right)x - 2a^2 k_0 \right]^2}{1 - i \frac{t_c}{t}} \right\} e^{i \frac{x^2}{4a^2} \frac{t_c}{t}}$$

$$\psi(x,t) = \frac{1}{(2\pi a^2)^{1/4}} \frac{1}{\sqrt{1 + i \frac{t_c}{t}}} \exp\left\{ -\frac{(t_c/t)^2}{4a^2} \frac{(x - v_g t)^2}{(1 - i \frac{t_c}{t})} \right\} e^{i \frac{x^2}{4a^2} \frac{t_c}{t}}$$

where  $v_g = \frac{\hbar k}{m} \Big|_{k_0} = \frac{\partial \omega(k)}{\partial k} \Big|_{k_0}$

Note:  $t_c$  appears just as expected from part (b)



(d) The probability density in position

$$P(x,t) = |\psi(x,t)|^2 = \left| \frac{1}{(2\pi a^2)^{1/4}} \frac{1}{\sqrt{1+i\frac{t}{t_c}}} \exp\left\{-\frac{(t_c/t)^2}{4a^2} \frac{(x-v_g t)^2}{(1-i\frac{t_c}{t})}\right\} e^{i\frac{x^2}{4a^2} \frac{t_c}{t}} \right|^2$$

$$= \frac{1}{\sqrt{2\pi} a^2} \left| \frac{1}{\sqrt{1+i\frac{t}{t_c}}} \right|^2 \left| \exp\left\{-\frac{(t_c/t)^2}{4a^2} \frac{(x-v_g t)^2}{(1-i\frac{t_c}{t})}\right\} \right|^2 \left| e^{i\frac{x^2}{4a^2} \frac{t_c}{t}} \right|^2$$

Aside:  $|e^{i\frac{x^2}{4a^2} \frac{t_c}{t}}|^2 = 1$  since  $|e^{i\phi}| = 1$  for any real  $\phi$

$$\left| \frac{1}{\sqrt{1+i\frac{t}{t_c}}} \right|^2 = \frac{1}{\sqrt{1+i\frac{t}{t_c}}} \frac{1}{\sqrt{1-i\frac{t}{t_c}}} = \frac{1}{\sqrt{1+\frac{t^2}{t_c^2}}}$$

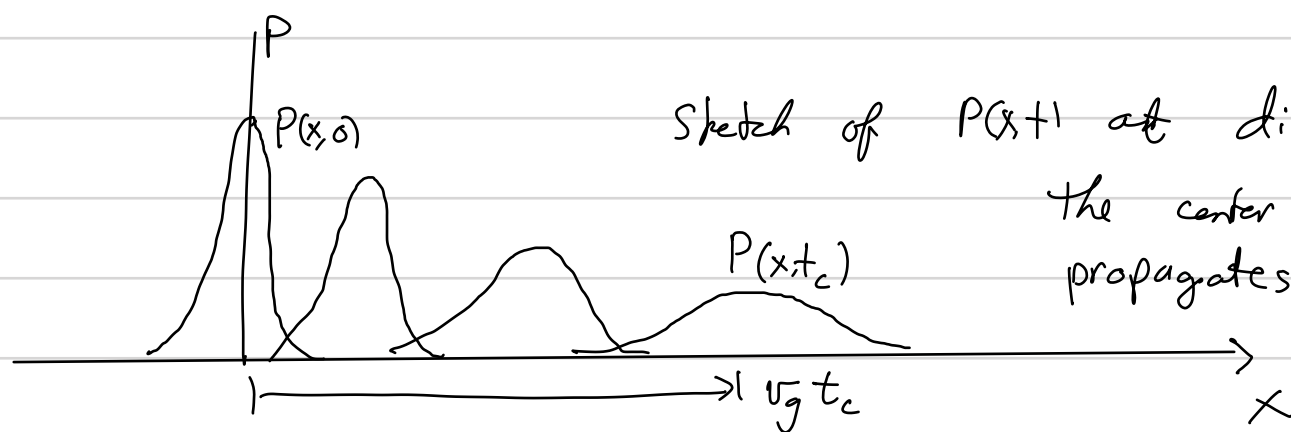
$$\left| \exp\left\{-\frac{(t_c/t)^2}{4a^2} \frac{(x-v_g t)^2}{(1-i\frac{t_c}{t})}\right\} \right|^2 = \exp\left\{-\frac{(t_c/t)^2}{4a^2} \frac{(x-v_g t)^2}{(1-i\frac{t_c}{t})}\right\} \exp\left\{-\frac{(t_c/t)^2}{4a^2} \frac{(x-v_g t)^2}{(1+i\frac{t_c}{t})}\right\}$$

$$= \exp\left\{-\frac{(t_c/t)^2}{4a^2} (x-v_g t) \left[ \frac{1}{1-i\frac{t_c}{t}} + \frac{1}{1+i\frac{t_c}{t}} \right]\right\} = \exp\left\{-\frac{(x-v_g t)^2}{2a^2 \left(1+\frac{t^2}{t_c^2}\right)}\right\}$$

Putting it all together,

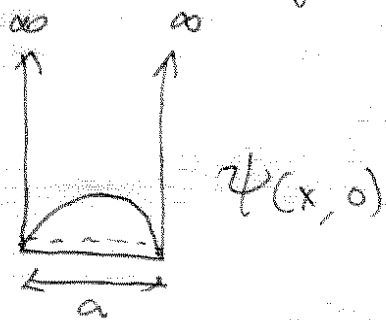
$$P(x,t) = \frac{1}{\sqrt{2\pi} a^2(t)} \exp\left\{-\frac{(x-v_g t)^2}{2a^2(t)}\right\}, \quad a^2(t) = a^2 \left(1 + \frac{t^2}{t_c^2}\right)$$

This is exactly what we guessed with basic estimation. The wave packet spreads, and doubles in width in a time  $t_c = \frac{2ma^2}{\hbar}$

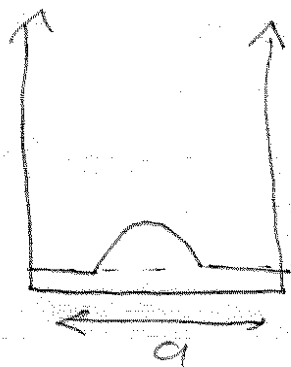


### Problem 3: The Sudden Approximation

At  $t=0$  a particle is prepared in the ground state of an infinite box



Suddenly, the box is expanded to double its size



Right after the expansion the wave function is the same.

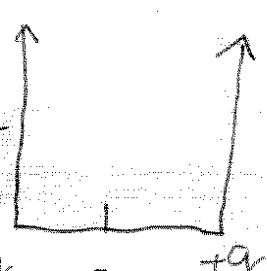
However,  $\psi(x,0_+)$  is no longer a stationary state.

thus it will evolve with time. The problem is to find  $\psi(x,t)$ .

Trick: From the symmetry of the problem, it makes more sense to take the origin at the center of the well

Before

$$u_n = \begin{cases} \sqrt{\frac{2}{a}} \cos\left(\frac{n\pi x}{a}\right) & n=1,3,5 \\ \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi x}{a}\right) & n=2,4,6 \end{cases}$$



After  $a \rightarrow 2a$

$$u'_n(x) = \begin{cases} \sqrt{\frac{1}{2a}} \cos\left(\frac{n\pi x}{2a}\right) & n=1,3,5 \\ \sqrt{\frac{1}{2a}} \sin\left(\frac{n\pi x}{2a}\right) & n=2,4,6 \end{cases}$$

(a) The wave function  $\psi(x, 0_+)$  can be expanded in terms of the new energy eigenfunctions  $\{u'_n(x)\}$

$$\psi(x, 0_+) = \sum_n c_n u'_n(x)$$

$$\text{where } \psi(x, 0_+) = \begin{cases} \sqrt{\frac{2}{a}} \cos\left(\frac{\pi x}{a}\right) & -\frac{a}{2} \leq x \leq \frac{a}{2} \\ 0 & \text{otherwise} \end{cases}$$

$$\text{And } c_n = \langle u'_n | \psi \rangle = \int_{-\infty}^{\infty} u'_n(x)^* \psi(x, 0_+) dx$$

$$= \begin{cases} \int_{-\frac{a}{2}}^{+\frac{a}{2}} dx \left( \frac{1}{\sqrt{a}} \cos\left(\frac{n\pi x}{2a}\right) \right) \left( \sqrt{\frac{2}{a}} \cos\left(\frac{\pi x}{a}\right) \right) & n = 1, 3, 5, \dots \\ \int_{-\frac{a}{2}}^{+\frac{a}{2}} dx \left( \frac{1}{\sqrt{a}} \sin\left(\frac{n\pi x}{2a}\right) \right) \left( \sqrt{\frac{2}{a}} \cos\left(\frac{\pi x}{a}\right) \right) & n = 2, 4, 6, \dots \end{cases}$$

Now, sine is an odd function over the interval whereas cosine is an even function

$$\Rightarrow \boxed{c_n = 0 \text{ for } n = 2, 4, 6, \dots}$$

For  $n = 1, 3, 5, \dots$

$$c_n = \frac{\sqrt{2}}{a} \int_{-\frac{a}{2}}^{+\frac{a}{2}} \cos\left(\frac{n\pi x}{2a}\right) \cos\left(\frac{\pi x}{a}\right) dx \quad (\text{next page})$$

Aside:  $\cos \theta_1, \cos \theta_2 = \frac{1}{2} [\cos(\theta_1 + \theta_2) + \cos(\theta_1 - \theta_2)]$

$\Rightarrow$  For  $n=1, 3, 5$

$$\begin{aligned}
 c_n &= \frac{\sqrt{2}}{a} \left( \frac{1}{2} \int_{-\frac{a}{2}}^{\frac{a}{2}} dx \left[ \cos\left(\frac{(n+2)\pi x}{2a}\right) + \cos\left(\frac{(n-2)\pi x}{2a}\right) \right] \right) \\
 &= \frac{\sqrt{2}}{\pi} \left[ \frac{1}{n+2} \sin\left(\frac{(n+2)\pi x}{2a}\right) + \frac{1}{n-2} \sin\left(\frac{(n-2)\pi x}{2a}\right) \right]_{-\frac{a}{2}}^{\frac{a}{2}} \\
 &= \frac{2\sqrt{2}}{\pi} \left[ \frac{1}{n+2} \sin\left(\frac{(n+2)\pi}{4}\right) + \frac{1}{n-2} \sin\left(\frac{(n-2)\pi}{4}\right) \right] \\
 &= \frac{2\sqrt{2}}{\pi} \left[ \frac{1}{n+2} \cos\left(\frac{n\pi}{4}\right) - \frac{1}{n-2} \cos\left(\frac{n\pi}{4}\right) \right] \\
 &= -\frac{8\sqrt{2}}{\pi} \frac{1}{n^2-4} \cos\left(\frac{n\pi}{4}\right)
 \end{aligned}$$

$$\begin{aligned}
 c_n &= -\frac{8\sqrt{2}}{\pi} \frac{1}{n^2-4} \cos\left(\frac{n\pi}{4}\right) \quad n=1, 3, 5, \dots \\
 &0 \quad n=2, 4, 6, \dots
 \end{aligned}$$

$$\psi(x, a_+) = \sum_{n \text{ odd}} c_n u_n'(x)$$

$$\text{nodal: } u_n'(x) = \begin{cases} \frac{1}{\sqrt{a}} \cos\left(\frac{n\pi x}{2a}\right) & -a \leq x < a \\ 0 & \text{otherwise} \end{cases}$$

(b) The possible energies that can be measured are the energy eigenvalues.

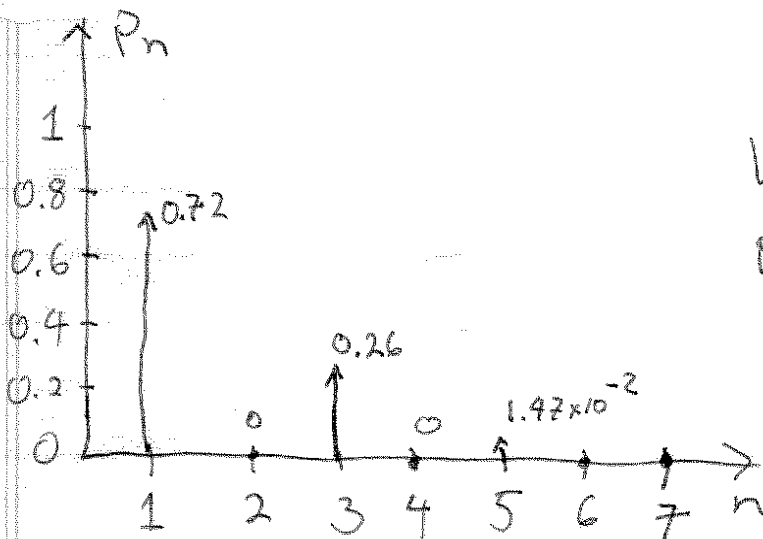
For an infinite box of width  $2a$  the eigenvalues are

$$E_n = n^2 \frac{\pi^2 \hbar^2}{2m(2a)^2} = n^2 \frac{\pi^2 \hbar^2}{8ma^2}$$

The ~~prob~~ probability of measuring  $n^{\text{th}}$  eigenvalue is the square of the probability amplitude

$$P_n = |c_n|^2 = \begin{cases} 0 & n = 2, 4, 6, \dots \\ \frac{64}{\pi^2} \frac{1}{(n^2-4)^2} & n = 1, 3, 5, \dots \end{cases}$$

Sketch



We see that the probability to be in the excited states falls us very rapidly with  $n$

$$P_n \sim \frac{1}{n^4} \quad n \gg 1$$

(c) Given the expansion of the wave function at  $t=0$  in terms of stationary states, we know a power series expansion for the solution at all times.

$$\Psi(x,t) = \sum_{n \text{ odd}} c_n \psi_n'(x) e^{-iE_n t/\hbar}$$

where  $c_n = (-1)^{\frac{n+1}{2}} \frac{8}{\pi} \frac{1}{n^2-4}$

$$\psi_n' = \begin{cases} \frac{1}{\sqrt{2}} \cos\left(\frac{n\pi x}{2a}\right) & -a \leq x \leq a \\ 0 & \text{otherwise} \end{cases}, \quad E_n = n^2 \frac{\pi^2 \hbar^2}{8ma^2}$$

(d) From our solution in part (b) we see that  $\sim 98\%$  of the probabilities is concentrated in the modes  $\psi_1'(x)$  and  $\psi_3'(x)$ . This makes sense intuitively.

$$\begin{aligned} \Rightarrow \Psi(x,t) &\approx c_1 \psi_1'(x) e^{-iE_1 t/\hbar} + c_3 \psi_3'(x) e^{-iE_3 t/\hbar} \\ &= e^{-iE_1 t/\hbar} \left( c_1 \psi_1'(x) + c_3 \psi_3'(x) e^{-i\Delta E_{31} t/\hbar} \right) \end{aligned}$$

where  $\Delta E_{31} \equiv E_3 - E_1$

(Next Page)

Now, as in free space, we expect the initially localized wave packet to spread. However, unlike free space, we don't expect this to go indefinitely. In fact, after a finite time, the packet will reconstitute itself because of the discrete nature of spectrum. Under the approximation that only  $n=1$  and  $n=3$  modes are present, the probability density

$|\Psi(x,t)|^2$  will oscillate at the period  $T \approx \frac{2\pi\hbar}{E_3 - E_1}$   
 or  $T \approx \frac{\pi\hbar}{2E_1}$

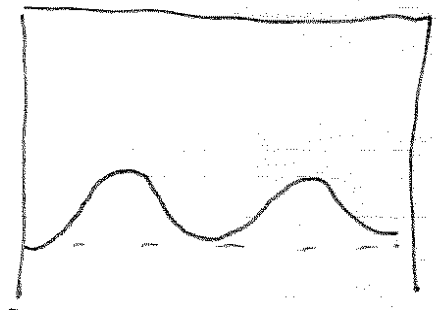
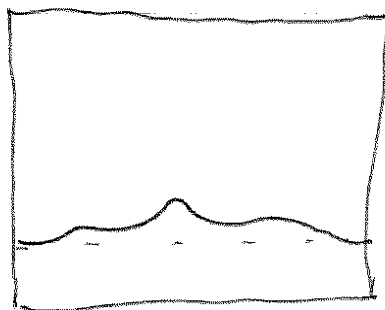
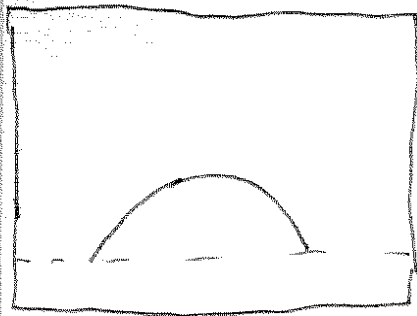
This is not exact, but a good approximation

$$t=0 \quad |\Psi(x,t)|^2 = |c_1 u_1'(x) + c_3 u_3'(x)|^2$$

$$t = \frac{T}{2} \quad |\Psi(x,t)|^2 = |c_1 u_1'(x) - c_3 u_3'(x)|^2$$

$$t = T \quad |\Psi(x,t)|^2 = |c_1 u_1'(x) + c_3 u_3'(x)|^2$$

Sketch (See Mathematica Notebook on web.)



$$t=0$$

$$t = \frac{T}{4} = \frac{\pi\hbar}{2E_3 - E_1}$$

$$t = \frac{T}{2} = \frac{\pi\hbar}{E_3 - E_1}$$

